

**Final Examination, 12 Dec 2002 (Solutions)**  
**SM311O (Fall 2002)**

1. (a) Let  $\mathbf{v} = \langle x^2z, \sin(ay), z\sqrt{x} \rangle$  where  $a$  is a constant. Determine  $a$  so that the divergence of  $\mathbf{v}$  vanishes at the point  $P = (4, 0, 1)$ .

**Solution:**  $\text{div } \mathbf{v} = 2xz + a \cos(ay) + \sqrt{x}$ . At  $P$  the divergence has the value  $10 + a$ , which vanishes when  $a = -10$ .

- (b) Let  $\mathbf{v} = \langle y^2 - x, y - x^2, 0 \rangle$ . Find the curl of  $\mathbf{v}$ . Is this flow irrotational anywhere?

- (c) Prove the identity  $\nabla \times \nabla \phi = \mathbf{0}$  if  $\phi$  is an arbitrary function of  $x$ ,  $y$ , and  $z$ .

**Solution:**  $\nabla \times \nabla \phi = \nabla \times \langle \phi_x, \phi_y, \phi_z \rangle = \langle \phi_{zy} - \phi_{yz}, \phi_{zx} - \phi_{xz}, \phi_{yx} - \phi_{xy} \rangle = \langle 0, 0, 0 \rangle$  because the order of differentiation does not matter for smooth (at least twice differentiable) functions.

2. Verify by direct differentiation if

- (a)  $u(z) = e^{2z} \cos 2z$  is a solution of  $u'' + au' + bu = 0$  for any pair  $(a, b)$ .

**Solutions:** With  $u = e^{2z} \cos 2z$  the differential operator  $u'' + au' + bu$  takes the form

$$e^{2z} ((2a + b) \cos(2z) - 2(4 + a) \sin(2z)).$$

. This expression vanishes for ALL  $z$  if  $2a + b = 0$  and  $a + 4 = 0$ , or if  $a = -4$  and  $b = 8$ .

- (b)  $u(x, y) = \sin 3x \cos 4y$  is an eigenfunction of the Laplace operator  $-\frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial y^2}$ . What is the eigenvalue?

**Solution:** A function  $u$  is an eigenfunction of  $-\frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial y^2}$  with eigenvalue  $\lambda$  if  $-\frac{\partial^2 u}{\partial x^2} - \frac{\partial^2 u}{\partial y^2} = \lambda u$ .

Let  $u = \sin 3x \cos 4y$ . Then  $-\frac{\partial^2 u}{\partial x^2} - \frac{\partial^2 u}{\partial y^2} = 25 \sin 3x \cos 4y = 25u$ . So  $u$  is an eigenfunction with eigenvalue 25.

3. (a) Give a parametrization for the plane that passes through the points  $(1, 1, 0)$ ,  $(0, 2, 2)$ , and  $(3, 0, 3)$ .

**Solution:** First we use the three points on the plane and find two vectors that are parallel with the plane:  $(1, 1, 0)$  and  $(0, 2, 2)$  give  $\mathbf{r}_1 = \langle -1, 1, 2 \rangle$ ;  $(0, 2, 2)$  and  $(3, 0, 3)$  give  $\mathbf{r}_2 = \langle 3, -2, 1 \rangle$ . Next we form the cross product of  $\mathbf{r}_1$  and  $\mathbf{r}_2$  to get  $\mathbf{n}$  a normal vector to the plane:

$$\mathbf{n} = \mathbf{r}_1 \times \mathbf{r}_2 = \langle 5, 7, -1 \rangle.$$

Let  $(x, y, z)$  be any point on the plane. Considering that  $(1, 1, 0)$  is also on the plane, the vector  $\mathbf{r}_3 = \langle x - 1, y - 1, z \rangle$  is parallel with the plane. Then  $\mathbf{n} \cdot \mathbf{r}_3 = 0$  i.e.  $5x + 7y - z = 12$ .

- (b) Find a unit normal vector to the surface of the upper hemisphere of the Earth at the point whose longitude and latitude are 45 and 60 degrees, respectively.

**Solution:** The upper hemisphere is parametrized as

$$\mathbf{r}(u, v) = R \langle \cos u \cos v, \sin u \cos v, \sin v \rangle$$

where  $R$  is the Earth's radius and  $u$  and  $v$  are longitude and latitude. Then  $\mathbf{n} = \mathbf{r}_u \times \mathbf{r}_v$  is normal to the hemisphere. Now  $\mathbf{r}_u \times \mathbf{r}_v = \langle R^2 \cos(u) \cos(v)^2, R^2 \cos(v)^2 \sin(u), R^2 \cos(v) \sin(v) \rangle$ . The

magnitude of this vector is  $R^2 \cos v$ . Dividing  $\mathbf{n}$  by its magnitude yields the desired unit vector  $\mathbf{N} = \langle \cos(u) \cos(v), \cos(v) \sin(u), \sin(v) \rangle$ . Finally evaluating at  $u = \frac{\pi}{4}$  and  $v = \frac{\pi}{3}$  gives  $\mathbf{N} = \langle \frac{1}{2\sqrt{2}}, \frac{1}{2\sqrt{2}}, \frac{\sqrt{3}}{2} \rangle$ .

4. (a) The function  $\phi(x, y) = ax^2y^2 - by^2 + ax - by$  is the potential for a velocity vector field  $\mathbf{v}$ . Determine all values of  $a$  and  $b$  so that the velocity of the particle located at  $(2, -1)$  is  $\langle 1, 2 \rangle$ .

**Solution:**  $\mathbf{v} = \nabla\phi = \langle 2axy^2 + a, 2ax^2y - 2by - b \rangle$ . Evaluating  $\mathbf{v}$  at  $(2, -1)$  yields  $\langle 5a, -8a + b \rangle$  which equals  $\langle 1, 2 \rangle$  if  $a = \frac{1}{2}$  and  $b = \frac{18}{5}$ .

- (b) The function  $\psi(x, y) = ax^2 + xy + by^2$  is the stream function of a velocity field  $\mathbf{v}$ . Find all  $a$  and  $b$  so that the velocity of the particle located at  $(1, -2)$  has magnitude  $\frac{1}{2}$ .

**Solution:**  $\mathbf{v} = \langle \frac{\partial\psi}{\partial y}, -\frac{\partial\psi}{\partial x} \rangle = \langle x + 2by, -2ax - y \rangle$ . Evaluating this vector at  $(1, -2)$  and setting its magnitude equal to  $\frac{1}{2}$  yields the equation  $(2 - 2a)^2 + (1 - 4b)^2 = \frac{1}{4}$  for the set of all  $(a, b)$ s.

5. (a) Consider the velocity field  $\mathbf{v} = (x^2z - x)\mathbf{k}$ . Determine the flux of this fluid through the following two surfaces:

- i. a disk of radius 1 in the  $xy$ -plane and centered at the origin.

**Solution:** First parametrize the disk:  $\mathbf{r}(u, v) = \langle u \cos v, u \sin v, 0 \rangle$ . Next compute  $\mathbf{r}_u \times \mathbf{r}_v = \langle 0, 0, u \rangle$ . Then  $\int \int_S \mathbf{v} \cdot d\mathbf{r} = \int_0^{2\pi} \int_0^1 \mathbf{r}|_S \cdot \mathbf{r}_u \times \mathbf{r}_v \, dudv = \int_0^{2\pi} \int_0^1 \langle 0, 0, -u \cos v \rangle \cdot \langle 0, 0, u \rangle \, dudv = 0$  so as much fluid is passing through  $S$  from below to above it as in the opposite direction.

- ii. a disk of radius 1 in the plane  $z = 3$  and centered at the origin.

**Solution:** First parametrize the disk:  $\mathbf{r}(u, v) = \langle u \cos v, u \sin v, 3 \rangle$ . Next compute  $\mathbf{r}_u \times \mathbf{r}_v = \langle 0, 0, u \rangle$ . Then  $\int \int_S \mathbf{v} \cdot d\mathbf{r} = \int_0^{2\pi} \int_0^1 \mathbf{r}|_S \cdot \mathbf{r}_u \times \mathbf{r}_v \, dudv = \int_0^{2\pi} \int_0^1 \langle 0, 0, -u \cos v + 3u^2 \cos^2 v \rangle \cdot \langle 0, 0, u \rangle \, dudv = \frac{3\pi}{4}$ .

- (b) Use the Stokes Theorem or compute the appropriate surface integral to determine the flux of vorticity of  $\mathbf{v} = x^2\mathbf{k}$  through the surface of the upper hemisphere of a sphere of radius 2 centered at the origin.

**Solution 1)** Direct computation:  $\omega = \nabla \times \mathbf{v} = \langle 0, -2x, 0 \rangle$ .  $\mathbf{r}(u, v) = \langle 2 \cos u \cos v, 2 \sin u \cos v, 2 \sin v \rangle$ .  $\mathbf{r}_u \times \mathbf{r}_v = \langle 4 \cos u \cos^2 v, 4 \cos^2 v \sin u, 2 \sin 2v \rangle$ . Then  $\int \int_S \omega \cdot d\mathbf{A} = \int_0^{2\pi} \int_0^{\frac{\pi}{2}} \langle 0, -4 \cos u \cos v, 0 \rangle \cdot \langle 4 \cos u \cos^2 v, 4 \cos^2 v \sin u, 2 \sin 2v \rangle \, dv du = 0$ .

**2)** Using the Stokes Theorem: We need to compute the line integral  $\int_C \mathbf{v} \cdot d\mathbf{r}$ .  $\mathbf{r} = \langle 2 \cos t, 2 \sin t, 0 \rangle$ . But  $\mathbf{v}|_C = \langle 0, 0, 4 \cos^2 t \rangle$  is orthogonal to  $\mathbf{r}'(t) = \langle -2 \sin t, 2 \cos t, 0 \rangle$ . So the line integral will vanish.

6. Consider the following wave equation initial-boundary value problem:

$$u_{tt} = 4u_{xx}, \quad u(0, t) = u(3, t) = 0, \quad u(x, 0) = x(3 - x), \quad u_t(x, 0) = 0.$$

- (a) Use separation of variables and find the solution to this problem. Clearly indicate the process of separation of variables and the Fourier Series method used in obtaining this solution.

**Solution:** From separation of variables we get that

$$u(x, t) = \sum_{n=1}^{\infty} \left( a_n \cos \frac{2n\pi t}{3} + b_n \sin \frac{2n\pi t}{3} \right) \sin \frac{n\pi x}{3}.$$

But  $u_t(x, 0) = 0$  which implies  $b_n = 0$  for all  $n$ .  $u(x, 0) = x(3 - x)$  so

$$a_n = \frac{(x(3 - x), \sin \frac{n\pi x}{3})}{(\sin \frac{n\pi x}{3}, \sin \frac{n\pi x}{3})} = \frac{2}{3} \left( \frac{54}{n^3 \pi^3} - \frac{54 \cos(n\pi)}{n^3 \pi^3} \right).$$

So

$$u(x, t) = \frac{2}{3} \sum_{n=1}^{\infty} \left( \frac{54}{n^3 \pi^3} - \frac{54 \cos(n\pi)}{n^3 \pi^3} \right) \cos \frac{2n\pi t}{3} \sin \frac{n\pi x}{3}.$$

- (b) Use the first nonzero term of the above solution and estimate  $u(3/2, 3/4)$ .

**Solution:** With  $n = 1$  the solution takes the form

$$u(x, t) = \frac{72}{\pi^2} \cos \frac{2\pi t}{3} \sin \frac{\pi x}{3}.$$

So  $u(3/2, 3/4) = 0$ .

- (c) Use the first nonzero term of the above solution and estimate how long it takes for the wave to go through one complete vibration.

**Solution:** The period of  $\cos \frac{2\pi t}{3}$  is  $2\pi / (2\pi/3) = 3$ .

7. Let  $\mathbf{v} = \langle x^2 + y^2, 2xy \rangle$  be the velocity field of a fluid. Compute the acceleration  $\mathbf{a}$  of this flow. Does  $\mathbf{a}$  have a potential  $p$ ? If yes, find it.

**Solution:**  $\mathbf{a} = \mathbf{v} \cdot \nabla \mathbf{v} = \langle 2(x^3 + 3xy^2), 2(3x^2y + y^3) \rangle$ . This vector has zero curl so there is a function  $p$  such that  $\nabla p = \langle 2(x^3 + 3xy^2), 2(3x^2y + y^3) \rangle$ . From

$$\frac{\partial p}{\partial x} = 2(x^3 + 3xy^2), \quad \frac{\partial p}{\partial y} = 2(3x^2y + y^3),$$

we get that  $p = \frac{x^4}{2} + 3x^2y^2 + \frac{y^4}{2}$ .

8. Let  $\boldsymbol{\Omega}$  stand for the angular velocity of our planet.

- (a) Noting that our planet rotates once every 24 hours, compute  $\boldsymbol{\Omega}$  where  $\boldsymbol{\Omega} = \langle 0, 0, \Omega \rangle$ . What are the units of  $\Omega$ ?

**Solution:**  $\Omega = \frac{2\pi \text{ radians}}{(24 \text{ hours})(60 \text{ minutes/hour})(60 \text{ seconds/minute})} = 0.00007 \text{ rad/s}$ .

- (b) Use this value of  $\boldsymbol{\Omega}$  and estimate the values in the centripetal acceleration  $\boldsymbol{\Omega} \times (\boldsymbol{\Omega} \times \mathbf{r})$  where  $\mathbf{r}$  is the position vector to a typical point on the surface of the Earth. Assume that the radius of the Earth is 6000 kilometers.

**Solution:** Let  $P = (x, y, z)$  be a point on the planet. Note that  $\boldsymbol{\Omega} = \langle 0, 0, \Omega \rangle$ . Then  $\mathbf{r} = \langle x, y, z \rangle$  and  $\boldsymbol{\Omega} \times (\boldsymbol{\Omega} \times \mathbf{r}) = -\Omega^2 \langle x, y, 0 \rangle$ .  $x$  or  $y$  take their largest values at the equator which could be as large as 6000 kilometers. So the nonzero components of the centripetal acceleration could be as large as  $0.00007^2 (\text{rad/s})^2 \times 6000000 \text{ meters} = 0.0294 \text{ m/s}^2$ , considerably smaller than  $9.8 \text{ m/s}^2$  from the acceleration of gravity.

9. Consider an incompressible fluid occupying the basin

$$D = \{(x, y, z) | 0 \leq z \leq H\}.$$

Let  $\mathbf{v} = \langle v_1, v_2, v_3 \rangle$  be the velocity field of a motion generated in  $D$ . Suppose that we have been able to determine that

$$v_1(x, y, z) = 3x^2y^2 - x, \quad v_2(x, y, z) = yz,$$

but have only succeeded in measuring  $v_3$  at the bottom of the basin and that this value is

$$v_3(x, y, 0) = x + y.$$

Determine  $v_3$  everywhere in  $D$ . (Hint: What does incompressibility mean **mathematically**?)

**Solution:** From the equation of incompressibility we have

$$\frac{\partial v_1}{\partial x} + \frac{\partial v_2}{\partial y} + \frac{\partial v_3}{\partial z} = 0$$

or

$$\frac{\partial v_3}{\partial z} = -\frac{\partial v_1}{\partial x} - \frac{\partial v_2}{\partial y}.$$

Substituting the values of  $v_1$  and  $v_2$  in the above relation yields

$$\frac{\partial v_3}{\partial z} = -2xy^2.$$

Integrating this result with respect to  $z$  from 0 to  $z$  and using the value of  $v_3$  at  $z = 0$  yields

$$v_3 = -2xy^2z + x + y.$$

**Solution:** Since the flow is incompressible,

10. A flow is called geostrophic if the velocity  $\mathbf{v} = \langle u(x, y), v(x, y) \rangle$  and the pressure gradient  $\nabla p$  are related by

$$(*) \quad -fv = -\frac{1}{\rho} \frac{\partial p}{\partial x}, \quad fu = -\frac{1}{\rho} \frac{\partial p}{\partial y},$$

where  $\rho$ , a constant, is the density of the fluid, and  $f$  is the coriolis parameter.

- (a) Assuming that  $f$  is constant, prove that the divergence of  $\mathbf{v}$  must vanish.

**Solution:** From the equations of motion we have

$$u = -\frac{1}{\rho f} \frac{\partial p}{\partial y}, \quad v = \frac{1}{\rho f} \frac{\partial p}{\partial x}.$$

Now  $\text{div } \mathbf{v} = \frac{\partial v_1}{\partial x} + \frac{\partial v_2}{\partial y}$  which is equal to

$$-\frac{1}{\rho f} \frac{\partial^2 p}{\partial x \partial y} + \frac{1}{\rho f} \frac{\partial^2 p}{\partial y \partial x} = 0$$

- (b) Prove that the particle paths of a geostrophic flow and its isobars coincide.

**Solution:** Note that  $\mathbf{v} \cdot \nabla p = -\frac{1}{\rho f} \frac{\partial p}{\partial y} \frac{\partial p}{\partial x} + \frac{1}{\rho f} \frac{\partial p}{\partial x} \frac{\partial p}{\partial y} = 0$ . So  $\mathbf{v}$  and  $\nabla p$  are orthogonal. Since  $\nabla p$  is orthogonal to isobars, and since  $\mathbf{v}$  is tangential to particle paths, particle paths and isobars coincide.

- (c) Consider a high pressure field in a geostrophic flow in the northern hemisphere (where  $f > 0$ ). By appealing to the equations in (\*) explain whether this high pressure field results in a clockwise or a counterclockwise motion.

**Solution:** Without loss of generality, assume that the high pressure occurs at the origin of the coordinate system. Let  $P$  be a point in the first quadrant. Then  $\nabla p$  at  $P$  points toward the origin because 0 is a maximum of  $p$ . Then  $\frac{\partial p}{\partial x} \leq 0$  and  $\frac{\partial p}{\partial y} \leq 0$  at  $P$  (draw a picture to convince yourself of this). Going back to the geostrophic equations,  $u \geq 0$  and  $v \leq 0$  at  $P$  which indicates that the motion is clockwise.